



## Dark Matter search with the BDX-MINI experiment

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On behalf of BDX Collaboration

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- Background
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# Dark Matter Problem

Astrophysical observations suggest existence of DM

- Information only from gravitational interaction
- ⇒ No clue on DM nature

Common assumption: **thermal origin of DM**

- DM we see comes from an epoch of thermodynamical equilibrium with SM
- constrain on available mass range
- strong constraint on viable DM → SM interaction



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Thermal DM



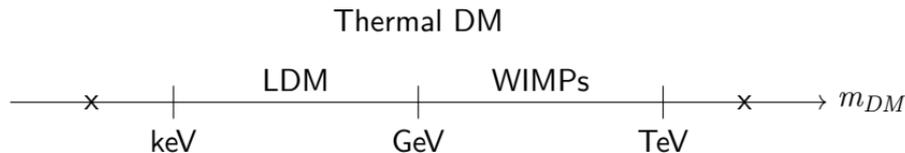
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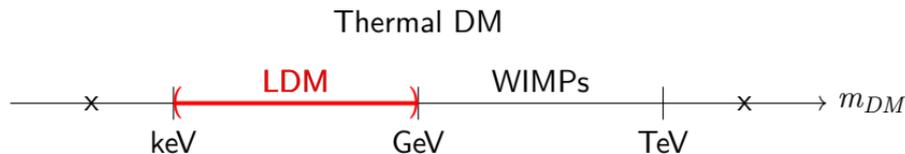
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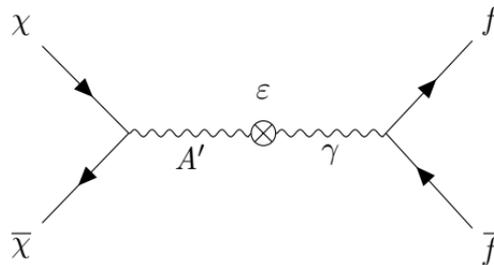
# Light Dark Matter - Dark Photon model

Simplest possibility: "vector portal"<sup>1</sup>

→  $U(1)$  gauge boson (**dark photon**) coupling to electric charge

$$\mathcal{L}_{LDM} \sim g_D A'_\mu J^\mu_\chi + \epsilon e A'_\mu J^\mu_{EM} + [\dots]$$

Annihilation in SM:



<sup>1</sup> For a comprehensive review: 1707.04591, 2005.01515, 2011.02157

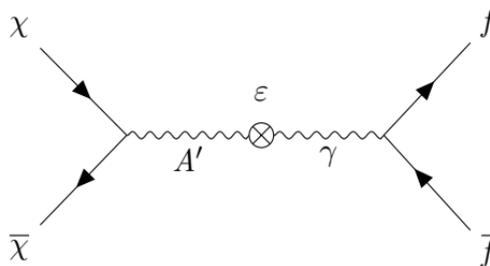
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Annihilation in SM:



Model parameters:

- Dark Photon mass  $m_{A'}$ , coupling to SM  $\epsilon$
- Dark Matter mass  $m_\chi$ , coupling to DM  $g_D$   
( $\alpha_D \equiv g_D^2/4\pi$ )

$$y \equiv \frac{g_D^2 \epsilon^2 e^2}{4\pi} \left( \frac{m_\chi}{m_{A'}} \right)^4 \sim \langle \sigma v \rangle_{\text{relic}} m_\chi^2$$

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# Light Dark Matter

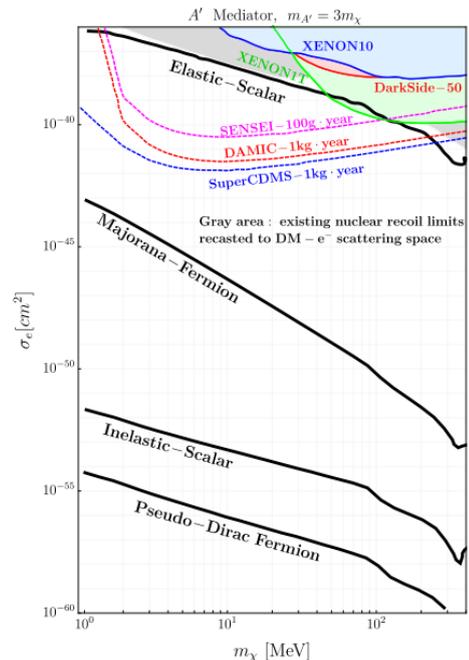
Direct detection not well suited for sub-GeV DM searches:

- DD experiments optimized for  $m_\chi > \text{GeV}$ 
  - $E_R \propto m_\chi^2 / m_N$
  - ⇒ very low recoil energy
- LDM-SM interaction cross section depends on impinging particle velocity
  - DD sensitivity strongly model-dependent

## LDM at accelerators

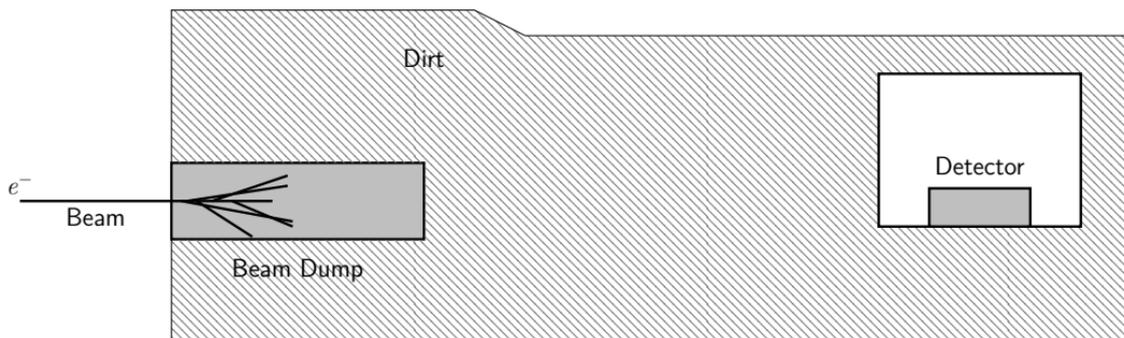
Accelerator based experiments at the *intensity frontier* uniquely suited to search for LDM:

- High intensity ⇒ increased possibility of DM production
- Production of relativistic DM ⇒ testing different models



# Beam Dump experiments

**Beam dump experiments:** direct detection of LDM produced by beam impinging on fixed target (beam dump)<sup>2</sup>

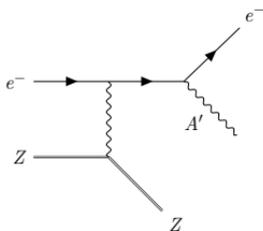
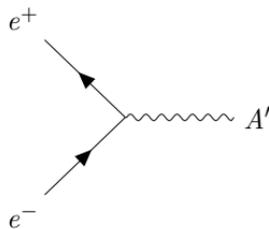
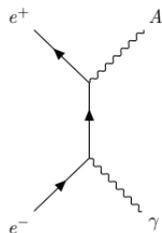
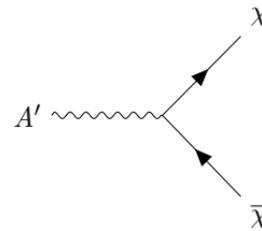


$\chi$  production

- $e^-$  beam impinging on target
- $\chi$  from decay of  $A'$  produced in the dump

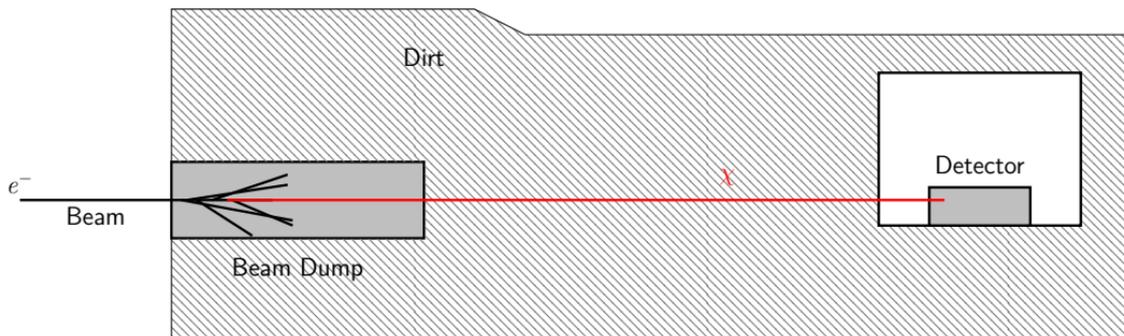
<sup>2</sup> Izaguirre et al., Phys. Rev. D 88, 114015

## Dark Photon model

Dark Matter production<sup>3</sup> $A'$ -strahlungResonant  $e^+e^-$   
annihilationNon-resonant  
 $e^+e^-$  annihilation $A'$  decay<sup>3</sup> L. Marsicano et al., Phys. Rev. Lett. 121, 041802

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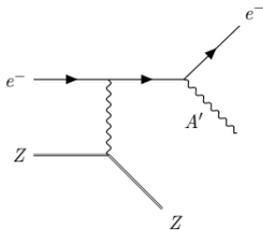
## $\chi$ interaction

- Detector placed behind the dump ( $\sim 10$  m)
- $\chi$  scattering through  $A'$  exchange

<sup>2</sup> Izaguirre et al., Phys. Rev. D 88, 114015

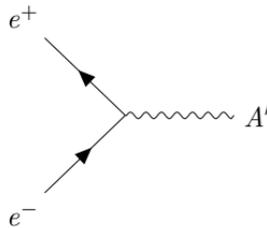
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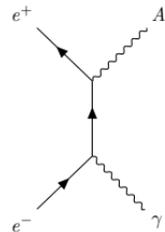


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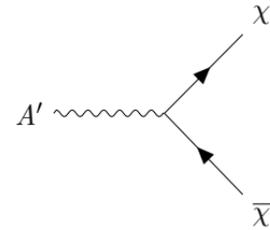
Resonant  $e^+e^-$  annihilation



Non-resonant  $e^+e^-$  annihilation

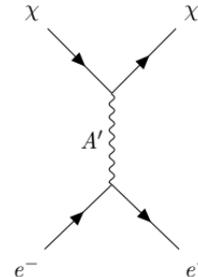


$A'$  decay

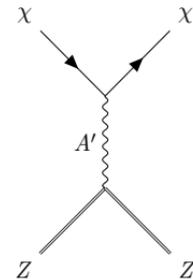


## Dark Matter interaction

Scattering on  $e^-$



Scattering on nuclei



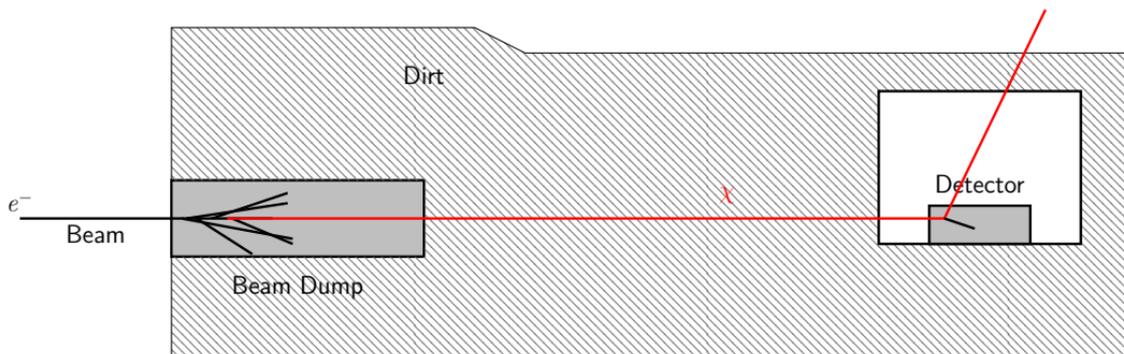
With  $\sim$  GeV DM recoil energy:

- $E_{rec}^e \sim 100$  MeV
- $E_{rec}^N \sim$  MeV

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# Beam Dump experiments

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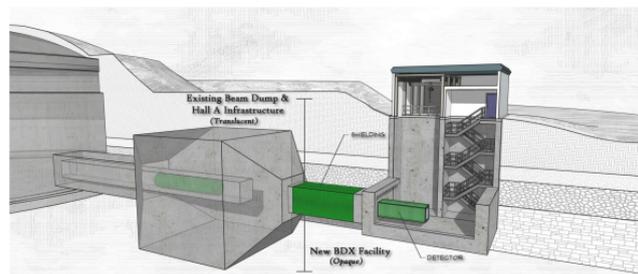
Number of signal events:  $S \propto \frac{\alpha_D \epsilon^4}{m_{A'}^4}$

<sup>2</sup> Izaguirre et al., Phys. Rev. D 88, 114015

## BDX

BDX is a **JLab experiment** approved by PAC46

- unique experiment able to produce and detect LDM
- beam dump experiment specifically aimed at LDM searches



Experiment designed with two goals:

### LDM production and detection

- Exploit CEBAF high-intensity beam
- Medium-high energy beam
- EM shower detection capability
- Fully parasitic

### Minimize background

- Shielding to filter beam-related background
- Multi layer veto for cosmogenic background
- Segmented detector
- Time resolution for detector-veto coincidence

# BDX - Experimental Setup

JLAB offers the best condition for BDX:

- High energy beam (11 GeV)
- High electron beam current ( $65 \mu\text{A}$ )
- Fully parasitic wrt Hall-A physic program (Moeller)

New facility to be built in front of Hall-A beam dump:

- new underground ( $\sim 8 \text{ m}$ ) hall
- 25 m downstream of Hall-A beam dump
- passive shielding ( $\sim 6.6 \text{ m}$  iron) to reduce beam related background
- $\sim 10 \text{ m}$  water equivalent overburden to reduce cosmogenic background



# BDX - Detector

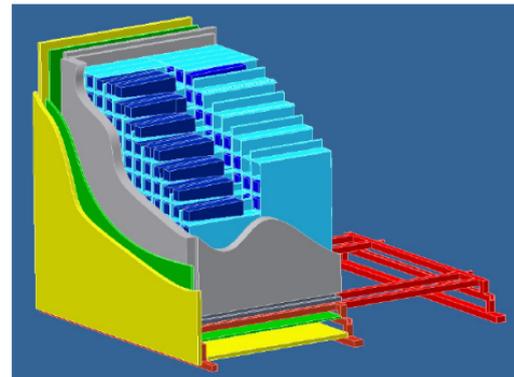
## Detector design

### Electromagnetic calorimeter:

- homogeneous calorimeter made with CsI(Tl) crystals read by SiPM

### Veto system:

- hermetic multi layer veto
- 2 layer of plastic scintillator counters read by WLS fibers and SiPM
- 5 cm lead vault between veto and calorimeter



### Modular detector arrangement:

- 1 module:  $10 \times 10$  crystals
    - $50 \times 50 \text{ cm}^2$  front face, 30 cm long
    - Module surrounded by veto
- total: 8 modules ( $\sim 2.6 \text{ m}$  length)

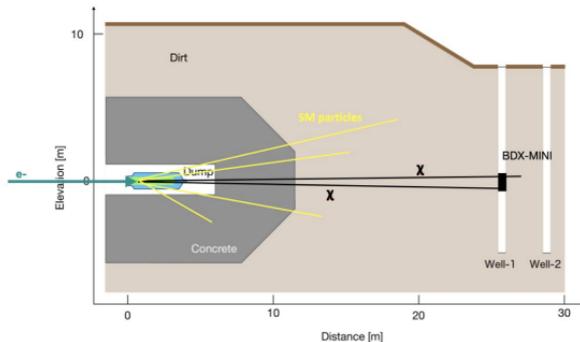
### Signal detection:

- EM shower ( $\gtrsim 100 \text{ MeV}$ ) and no corresponding activity in the active veto
- Signal efficiency  $\sim 20\%$

# BDX-MINI experiment

Pilot version of BDX:

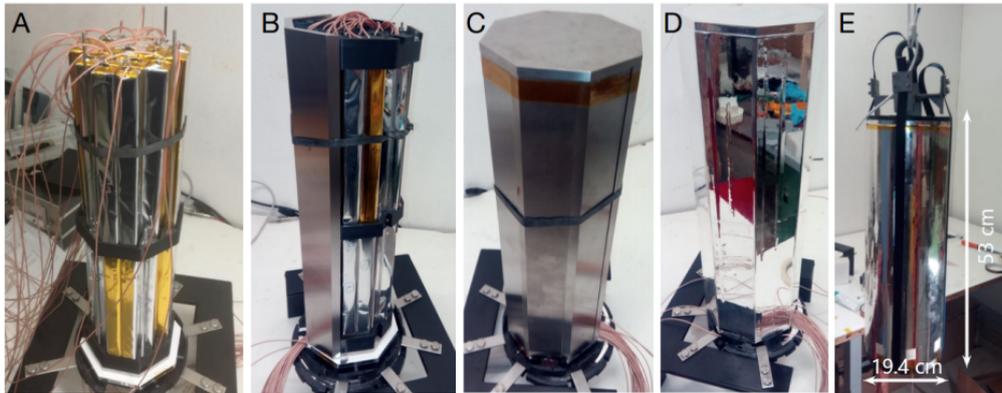
- 2.56 GeV  $e^-$  beam (10 GeV beam used for calibration)
- current up to 150  $\mu\text{A}$
- measurement alternating beam on and beam off data (beam on time  $\sim 50\%$ )
- accumulated  $2.54 \times 10^{21}$  EOT
- beam off measurements for cosmic background characterization



# BDX-MINI detector

## Electromagnetic calorimeter (ECal):

- 44 PbWO<sub>4</sub> crystals ( $4 \times 10^{-3} \text{ m}^3$  active volume)
- SiPM readout



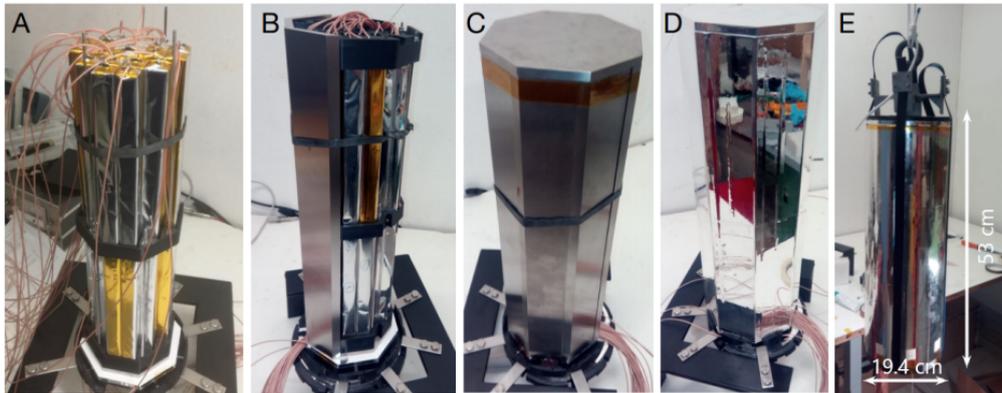
# BDX-MINI detector

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- 44  $\text{PbWO}_4$  crystals ( $4 \times 10^{-3} \text{ m}^3$  active volume)
- SiPM readout

## Veto system

- Active veto:
  - Octagonal (IV) and cylindrical (OV) plastic scintillator
  - Optically continuous
  - SiPM readout + WLS fibers light collection
- Passive tungsten shielding
  - 0.8 cm thick



# Detector performance

## Detector performance:

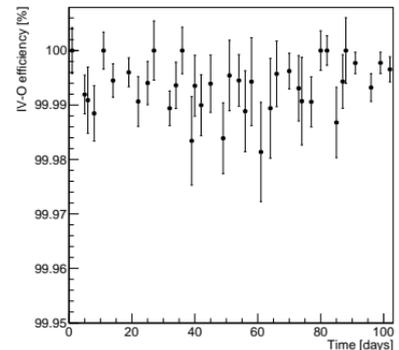
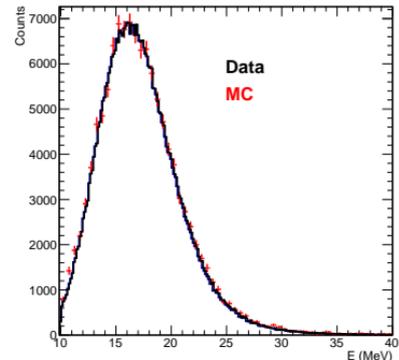
- Calibration with special 10 GeV run
- Stability monitored with cosmic  $\mu$

### ECal energy response

- Energy calibration determined from 10 GeV data
  - compared secondary  $\mu$  between data and MC
- Stability monitored with beam off data
  - Selected penetrating  $\mu$  with Landau peak

### VETO stability

- Veto efficiency monitored with cosmic muons
  - *tag-and-probe* method
  - trajectories selected with ECal energy deposition



# Backgrounds

Two main sources of background:

## Beam related background

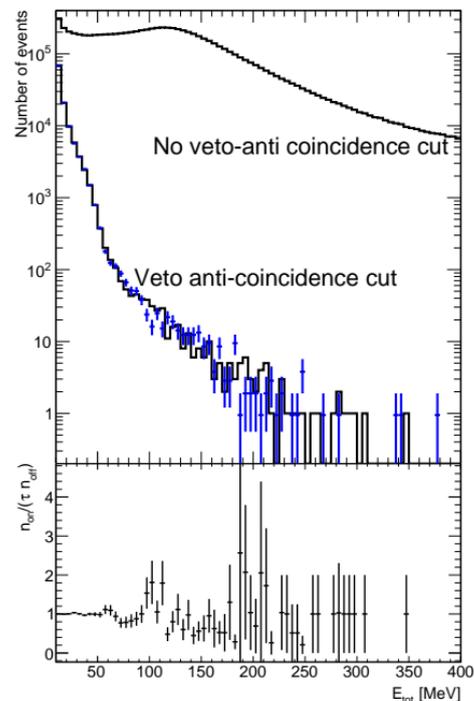
Yield estimate through MC simulations  
(FLUKA+GENIE+Geant4)

- MC simulation validated with in-situ measurement
- $\nu$  only background  $\rightarrow$  negligible:  $5.8 \times 10^{-23}$   $\nu/EOT$

## Cosmogenic background

- Continuous measurement  $\Rightarrow$  no rejection
- Charged particles rejected requiring veto anti-coincidence
- Further suppression can be achieved using energy cut

## Measured cosmic background



# Data analysis

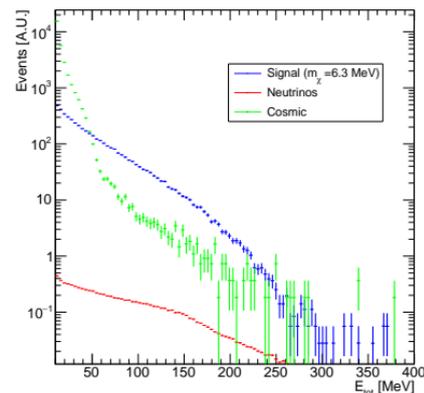
## Approach

**Blind analysis:** experiment sensitivity optimized with MC simulations and beam-off data

**Model:** ON-OFF problem

$$\mathcal{L} = \prod_j \left[ P(n_{\text{on}}^j; \mu_c^j + \mu_b^j + \alpha^j \cdot S) \cdot P(n_{\text{off}}^j; \mu_c^j \cdot \tau) \right]$$

- Data binned according to total energy deposition in ECal
- $n_{\text{on}}^j, n_{\text{off}}^j$ : measured number of events during beam-on/beam-off intervals ( $\tau = T_{\text{off}}/T_{\text{on}}$ )
- $\mu_c^j/\mu_b^j$ : expected number of cosmogenic/beam-related backgrounds events
- $\mu_b^j$  evaluated via MC,  $\mu_c^j$  treated as nuisance parameter.



**Systematic uncertainties:** described via ancillary pseudo-measurement factors in  $\mathcal{L}$  with Gaussian constraint

→ one-sided profile-likelihood test statistics to evaluate upper limit on  $S$

# Sensitivity optimization

**Idea:** improve sensitivity considering effect of data analysis cuts on background minimization and signal maximization

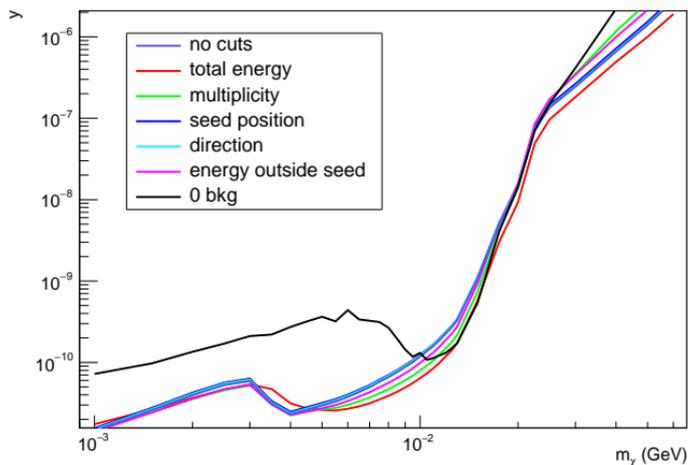
- beam-on data extrapolated from beam-off data
- used only events in anti-coincidence with veto
- evaluated upper limit on signal yield with different cuts and converted to exclusion limit

Tested different cuts

## Best condition

- Anti-coincidence with veto
- $E_{tot} > 40$  MeV
- Data split into 8 energy bins

⇒ Sensitivity highly enhanced



# Unblinding

**Last step:** unblinding and analysis of beam on data

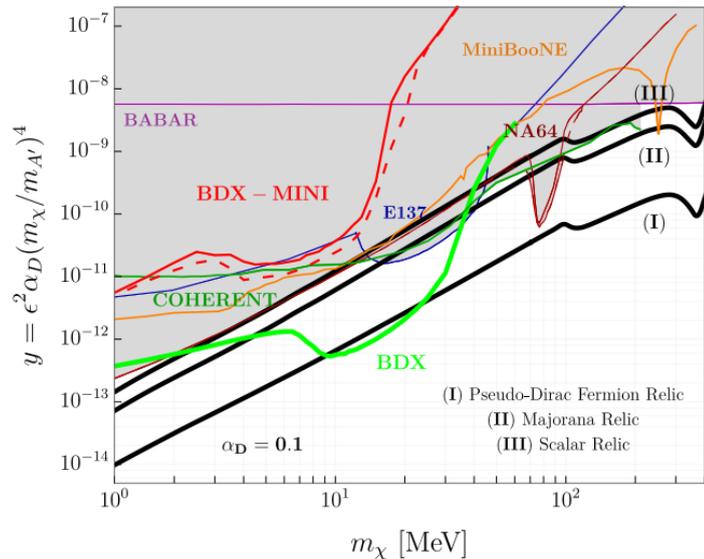
## Experimental results

Yields (for  $N_{EOT} = 2.54 \cdot 10^{21}$ )

- $N_{on} = 3623$
- $N_{off} = 3822$  ( $\tau = 1.054$ )

No excess is observed

→ evaluated 90% exclusion limit in the LDM parameter space



# Conclusions

- MeV-GeV Dark Matter range largely unexplored
- **Beam Dump eXperiment** at JLab: search for Dark Sector particles in the MeV-GeV mass range
- **BDX-MINI**: pilot version of BDX
  - First modern beam dump experiment searching for Light Dark Matter
  - Detector optimized for LDM searches
  - Analysis aimed to Light Dark Matter detection
  - Evaluated exclusion limit → competitive to flagship experiments
- Beam dump experiment at  $e$  beam dump highly sensitive to Light Dark Matter in the MeV-GeV range