Arc BLM activity before/after seed laser swap Jay Benesch

During the FY20 physics run, regions in the higher arcs have been activated in a manner never before seen. Chopper viewer images suggested a tail on the C beam was a cause of the difficulty. Accordingly, on March 19 while Hall C was down to switch experiments from a1n to d2n, the laser seeds for B and C were swapped. The B beam is defined by a narrow chopper slit and is low current so it it less likely than C beam, which passes through an open slit, to contain a tail. Response of selected BLMs in arcs 1 and 2 as a function of A and C current were compared for three six hour periods:

- A. March 10 1400-2000 with both high current halls, before the laser seed swap
- B. March 20 0700-1300 with Hall A the only high current hall
- C. March 23 0100-0700 with both high current halls, after the laser seed swap

Arc 1 BLMs 04, 09, 16, 22, 29, 33 and 37 were downloaded for the six hour periods. Arc 2 BLMs 4, 10, 16, 23, 29, 32 and 36 were also downloaded. The less correlated BLMs are shown first and ignored thereafter.



Figure 1. Arc 1 BLMs with no correlation to total current in arc. Period A, before seed change.



Figure 2. Arc 2 BLMs with little or no correlation to Hall C current. Period A, before seed change.



Figure 3. ILM1A09 vs Hall A and Hall C currents with both in the machine, before seed change. Correlation with Hall C (3H00) current is better than with Hall A, even though Hall A has five times the current for most of the period.



Figure 4. Same as figure 3 except points with both A and C currents under 1 μ A are removed. Cutting out the near-zero currents reduced R² for Hall A current by just over half while it reduced R² for Hall C current by only a fourth. This suggests that Hall C beam was responsible for the majority of the BLM activity even in arc 1 where Hall A current is five times as great.

I will now proceed to waste space by showing similar plots for the other four arc 1 BLMs whose data for period A I have on hand.



Figure 6. ILM1A29 has much better correlation with current



ILM1A33B = 25.317656 + 0.106908*IBC1H04CRCUR2

Summary of Fit	
RSquare	0.218446
RSquare Adj	0.218386
Root Mean Square Error	11.73196
Mean of Response	37.83817
Observations (or Sum Wgts)	12958
Parameter Estimates	

Term	Estimate	Std Error	t Ratio	Prob> t
Intercept	25.317656	0.23219	109.04	<.0001*
IBC1H04CRCUR2	0.106908	0.001777	60.18	<.0001*

Figure 7. ILM1A33 vs currents



Linear Fit

M1A37B = 41.74789	+0.174102	7*IBC1I	H04CF	RCUR2	
Summary of Fit					
RSquare	0	.18750	6		
RSquare Adj	0	0.187443			
Root Mean Square E	Error 2	1.0262	2		
Mean of Response	6	62.13791			
Observations (or Su	m Wgts)	1295	8		
Parameter Estin	nates				
Term	Estimat	e Std	Error	t Ratio	Prob> t
Intercept	41.7478	9 0.41	16134	100.32	<.0001*
IBC1H04CRCUR2	0.174102	7 0.00)3184	54.68	<.0001*





-----Linear Fit

Linear Fit

ILM1A33B = 9.103417 + 1.1245053*IBC3H00CRCUR4

Summary of Fit	
RSquare	0.581643
RSquare Adj	0.58161
Root Mean Square Error	8.583504
Mean of Response	37.83817
Observations (or Sum Wgts)	12958

Parameter Estimates

Term	Estimate	Std Error	t Ratio	Prob> t
Intercept	9.103417	0.226991	40.10	<.0001*
IBC3H00CRCUR4	1.1245053	0.008379	134.21	<.0001*

140 120 100 40 20 0 0 10 20 0 0 10 20 30 IBC3H00CRCUR4

Bivariate Fit of ILM1A37B By IBC3H00CRCUR4

----Linear Fit

Linear Fit

IL	M1A37B = 15.49094 + 1.82548	315*IBC3H00CRCUF	{4
	Summary of Fit		
	RSquare	0.496098	
	RSquare Adj	0.49606	
	Root Mean Square Error	16.55861	
	Mean of Response	62.13791	
	Observations (or Sum Wots)	12958	

Observations (or Sum Wgts)

Term	Estimate	Std Error	t Ratio	Prob> t		
Intercept	15.49094	0.437892	35.38	<.0001*		
IBC3H00CRCUR4	1.8254815	0.016163	112.94	<.0001*		



I conclude that before the laser seed swap that the Hall C beam was the cause of perhaps two-thirds the arc 1 BLM activity. I will now examine period B when Hall C was off for ILMs 9, 29, 33 and 37.

Figure 10. BLMs 33 and 37 with only Hall A, B and D beams in the machine



Figure 12. BLM 1A29 after seed swap with four halls operating.



Figure 14. BLM 1A37 after seed swap with four halls operating.



Figure 16. Arc 2 BLMs with Hall $C > 1 \mu A$ before seed change.



Figure 18. Arc 2 BLMs after seed change with Hall $C > 1 \mu A$.

Unfortunately, if I evaluate the linear fits at 30 uA I find that there is modestly less BLM activity attributable to Hall C beam in arc 1 and no significant change in arc 2. Compare the last two columns in the table below. Scale of BLM response is very different in Arc 1 vs Arc 2.

able 1. Linear in coefficients and values at 50 uA before and after the laser seeu change.						
BLM	before_intercept	before_slope	after_intercept	after_slope	before@30uA	after@30uA
1A09	30.83	3.85	55.86	2.61	146	134
1A29	16.42	2.39	21.59	1.71	88	73
1A33	9.10	1.12	12.61	0.86	43	38
1A37	15.49	1.83	20.78	1.78	70	74
2A04	-29.83	63.77	-32.94	61.49	1883	1812
2A10	-731.49	208.94	-939.00	217.92	5537	5599
2A23	-611.52	185.66	-663.31	187.41	4958	4959
2A29	-349.11	99.01	-96.55	154.10	2621	4526

Table 1. Linear fit coefficients and values at 30 uA before and after the laser seed change.

Conclusion: none



Figure 19. Distributions of currents in period A, the "before" sample

Figure 20. Distribution of currents in period C, the "after" sample.

This data is provided the readers so they may decide whether to ask the author to make additional cuts on the data. Recall that these sets have both A and C currents $\ge 1 \ \mu$ A. Still no conclusion.